Issues (1. Evolving Environments)

- Is halo-mass really KING, and does it determine all the environmental dependence of galaxies?
- It is true that there are more massive galaxies in clusters due to accelerated mass assembly.
- Does the way they are assembled also matter? ...such as mergers or smoother accretion, which is likely to be environmentally dependent at high-z?
- Is the quenching of SF self-regulated or driven by environmental effects?
 - AGN/SB feedback (high mass) versus
 RP/tidal stripping (low mass)?

Issues (1. Evolving Environments)

- Are SF/AGN activities boosted in some galaxies in protoclusters?
 - How? Mergers/interactions? The mode?

- Are there any environmental effects on gaseous inflow/outflow processes?
- Inflow (cold streams): can be different between common haloes and isolated haloes?
 - Outflow: suppressed by IGM pressure?

Issues (1. Evolving Environments)

 How can we properly compare ancestors with descendants if their environments and halo masses change with time due to cluster/group growth through assembly?

- Galaxies were not always born or modified where they are now.
- Proto-clusters are much rarer than present-day clusters. Therefore, many progenitors of today's cluster galaxies were "not" in protoclusters.
 - Pre-processing can be important.

Discussion items and my preferable answers

1. Halo-mass is really king? Not sure yet. The way of mass assembly may also matter at high-z? 2. Quenching also by external effects? **Boost of SF/AGN activities in protoclusters?** Maybe yes in some galaxies, probably due to mergers, but a larger sample and IFU obs. are required. 4. Inflows/outflows depend on environment? Yes, these are the key mechanisms on top of the merger, and should be investigated by CE and IFU obs. 5. Ancestors/descendants relation in growing environments? Yes, this complicates our understanding of

environmental effects, and should be properly incorporated.